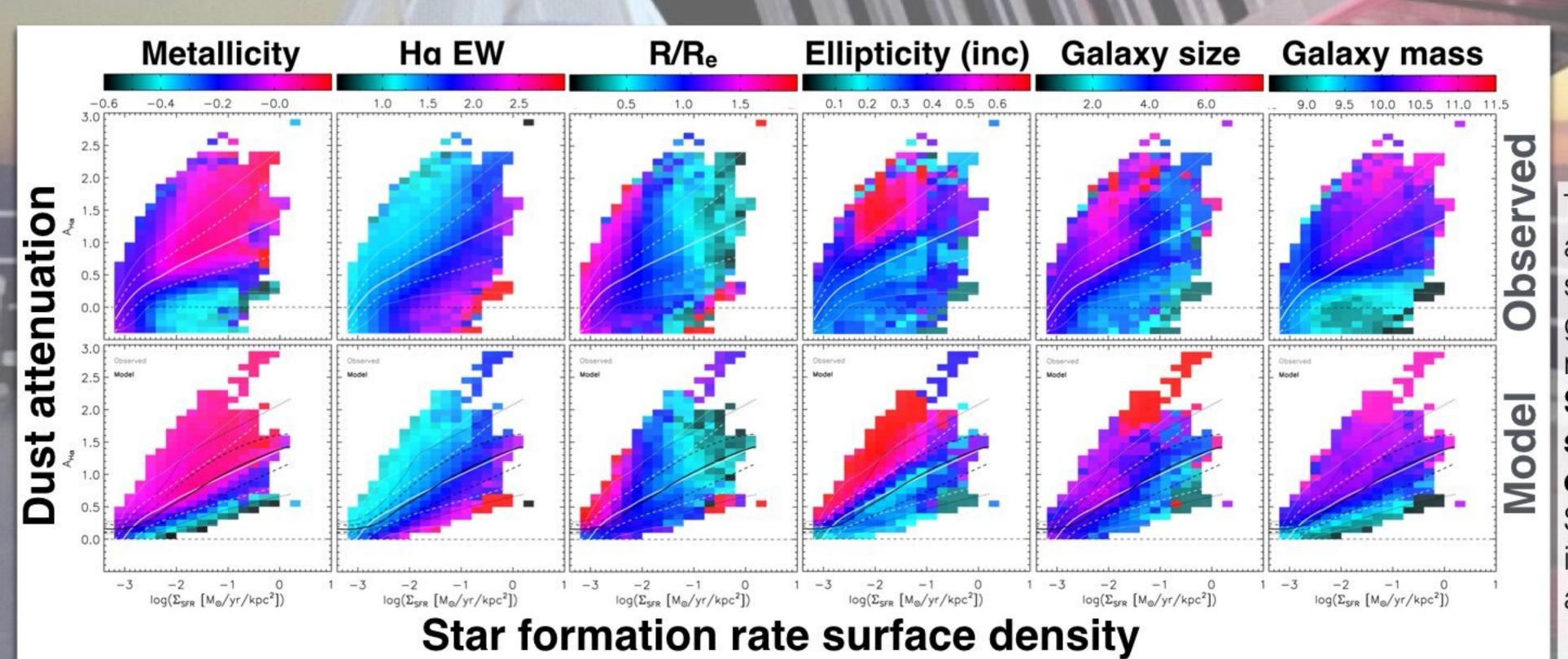
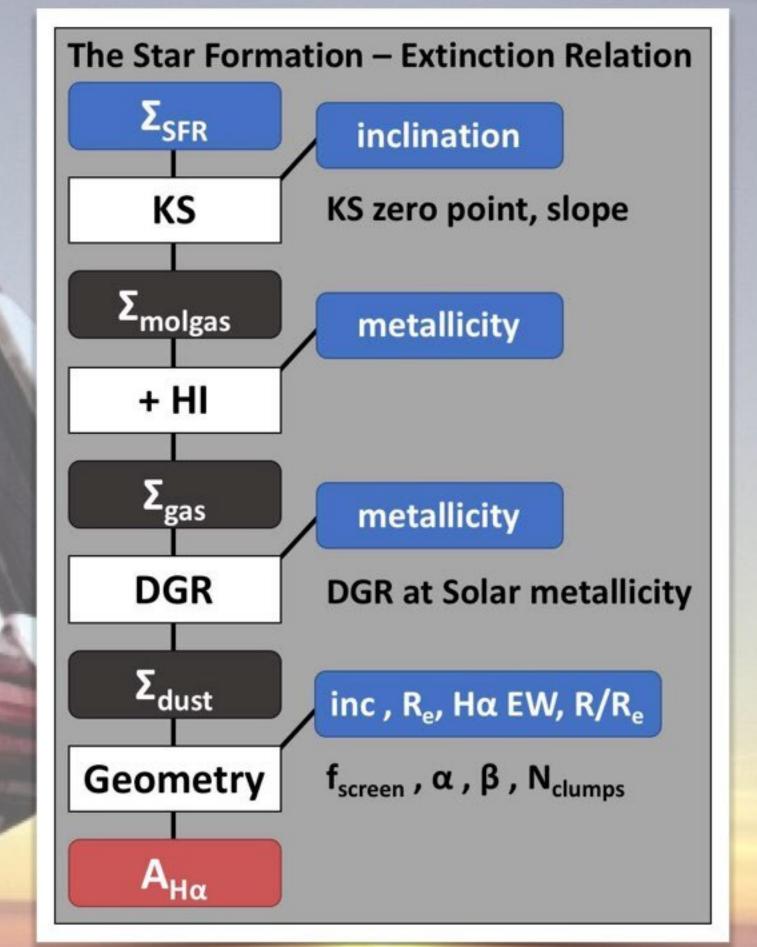
Resolving star formation across cosmic time

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MaNGA

The star formation - extinction relation within nearby galaxies





The relation between local star formation activity and dust attenuation, as well as its secondary dependencies on local and global galaxy properties, can be reproduced by a superlinear Kennicutt-Schmidt relation, a dust-to-gas ratio scaling linearly with metallicity and a clumpy dust geometry.

Simpler dust geometries fail, and inclination dependencies need to be accounted for.

Li, Wuyts+2019

Z~1 3D-HST vs IllustrisTNG

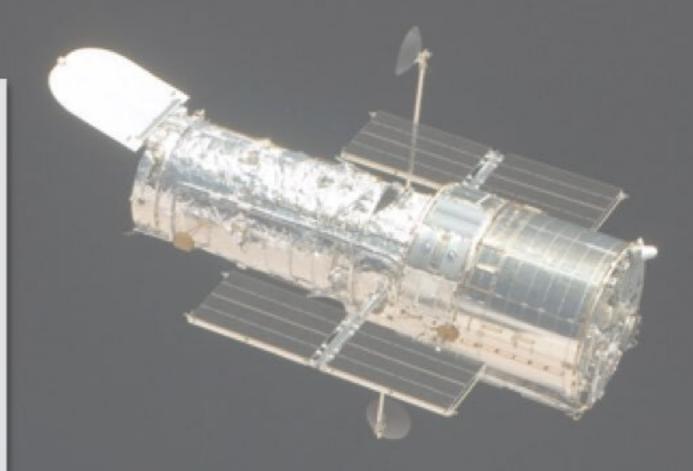
HST grism observations allow the extraction of Hα maps as a tracer of instantaneous star formation on kpc scales within star-forming galaxies (SFGs) at 0.7<z<1.5.

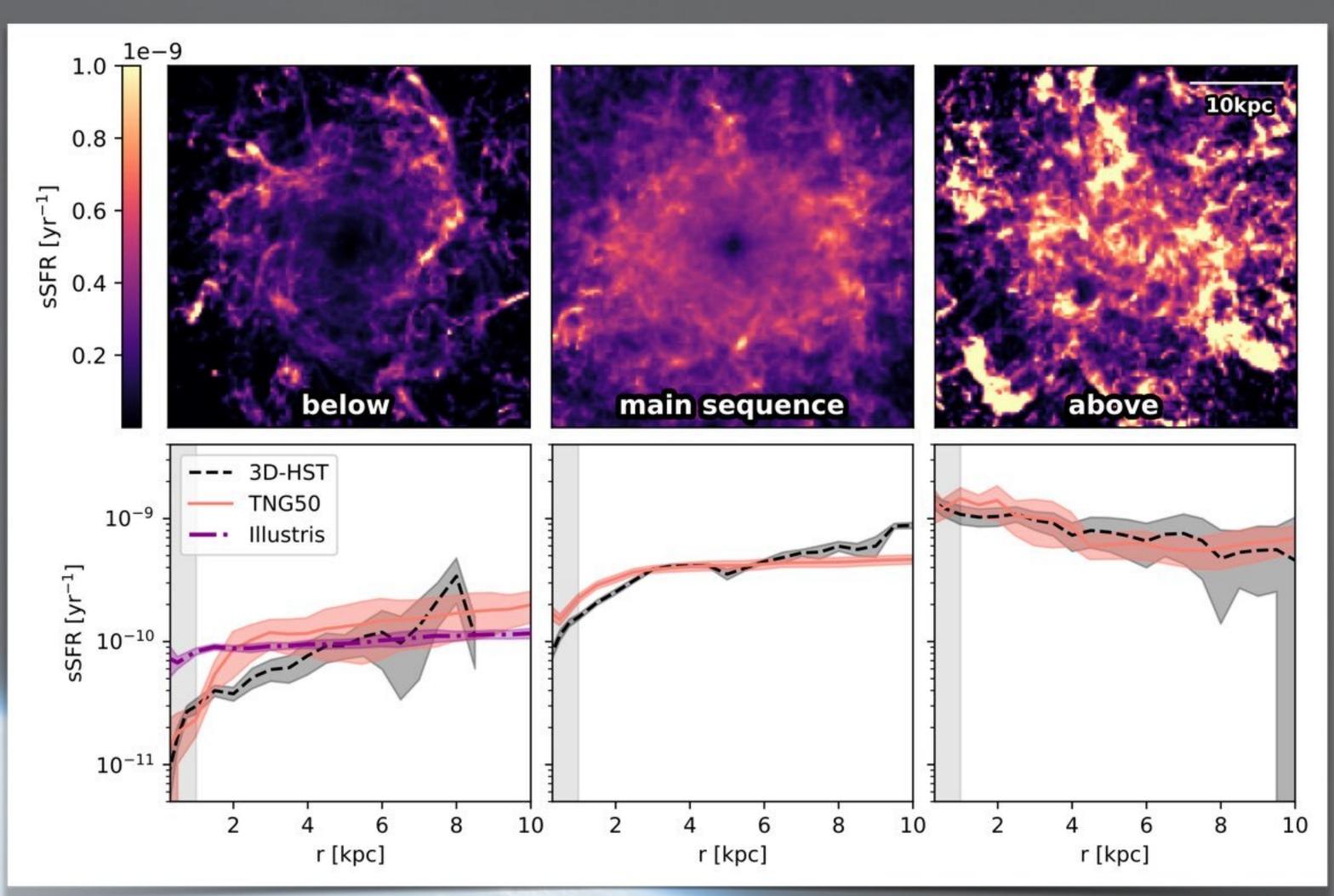
Dividing SFGs with stellar mass $log(M^*) = 10.5 - 11$ into those lying above/on/below the star-forming main sequence (MS), we find an overall shift in normalization of their specific star formation rate (sSFR = Σ_{SFR}/Σ_*) profiles. I.e., above-MS galaxies feature excess star formation at all radii.

The largest differences in sSFR are seen in the galaxy centres, where **below-MS SFGs** reveal the onset of **inside-out quenching**.

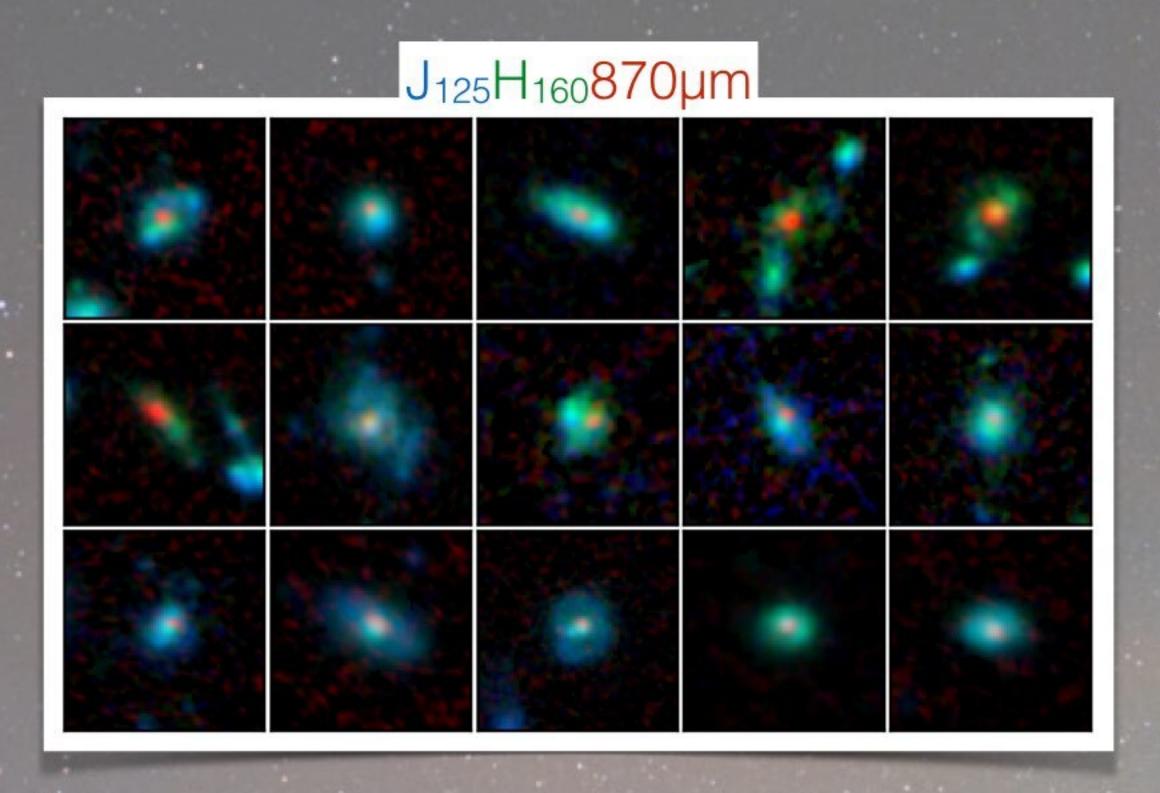
The **TNG50** simulation reproduces these trends remarkably well. Critically, this is no longer the case in a (smaller box) 'TNG variation' run where the **AGN kinetic** wind model is switched off.

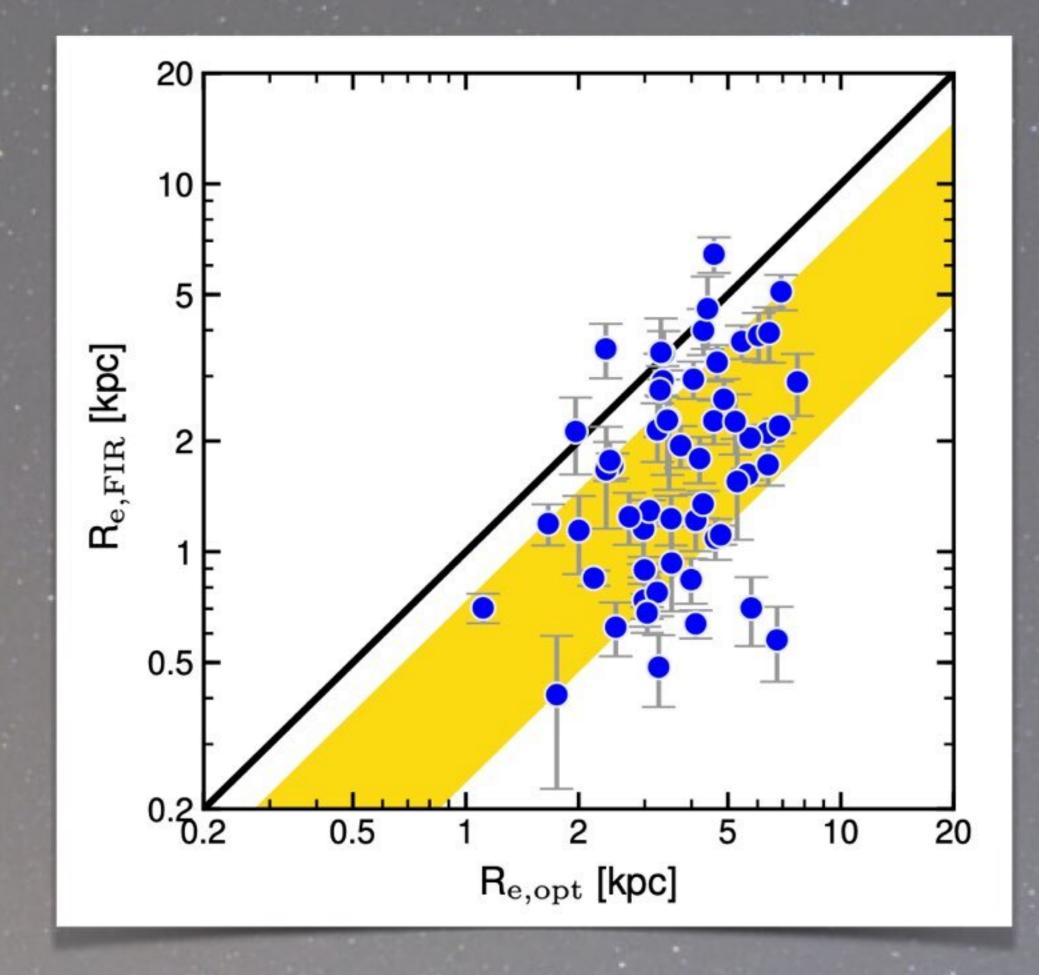
Nelson+2021

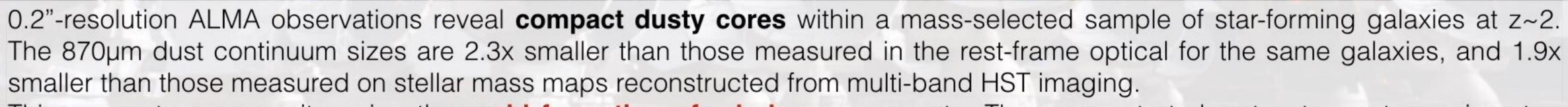




Z~2 log(M*)>11 Far-IR dust continuum







This suggests we are witnessing the **rapid formation of a bulge** component. The compact starburst puts most massive star-forming galaxies on the mass-size relation for quiescent $z\sim2$ galaxies within 300Myr, if the current star formation activity and its spatial distribution are maintained.

Tadaki+2017, 2020

